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Laschamp event 75ka (NAPIS −**75) and the duration of the North Atlantic palaeointensity stack since**

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Juerg Beer Carlo La j, Catherine Kissel, Alain Mazaud, James E. T. Channell and

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North Atlantic palaeointensity stack since 75 ka
(NAPIS-75) and the duration Exteed that the duration (NAPIS-75) and the duration
of the Laschamp event (NAPIS-75) and the duration
of the Laschamp event

of the Laschamp event
BY CARLO LAJ¹, CATHERINE KISSEL¹, ALAIN MAZAUD¹, RLO LAJ^1 , CATHERINE KISSEL¹, ALAIN MAZAUD¹,
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Six relative palaeointensity records from the north Atlantic Ocean were stacked
together to produce a new record for the last 75 kyr (NAPIS-75). Five of these Six relative palaeointensity records from the north Atlantic Ocean were stacked
together to produce a new record for the last 75 kyr (NAPIS-75). Five of these
records have been previously correlated at millennial scale and Six relative palaeointensity records from the north Atlantic Ocean were stacked
together to produce a new record for the last 75 kyr (NAPIS-75). Five of these
records have been previously correlated at millennial scale and together to produce a new record for the last 75 kyr (NAPIS-75). Five of these
records have been previously correlated at millennial scale and placed on the GISP2 records have been previously correlated at millennial scale and placed on the GISP2 age scale, the sixth record was tied to the others using magnetic susceptibility. From 75 ka the field strength exhibits some oscillation age scale, the sixth record was tied to the others using magnetic susceptibility. From
75 ka the field strength exhibits some oscillations, with a first minimum *ca*. 65 ka,
followed by a progressive increase to a broad m 75 ka the field strength exhibits some oscillations, with a first minimum $ca. 65$ ka, followed by a progressive increase to a broad maximum centred at $ca. 48$ ka. There is then a well-marked low at 40 ka, corresponding to followed by a progressive increase to a broad maximum centred at *ca*. 48 ka. There is then a well-marked low at 40 ka, corresponding to the directional anomaly of the Laschamp event. Another intensity low, observed at ca 34 ka, corresponds in age to the Mono Lake event. After a high at 33 ka and two lo Laschamp event. Another intensity low, observed at $ca.34$ ka, corresponds in age to the Mono Lake event. After a high at 33 ka and two lows at 30 and 24 ka with a broad maximum between, the field strength seems to slowly the Mono Lake event. After a high at 33 ka and two lows at 30 and 24 ka with a broad
maximum between, the field strength seems to slowly increase to the upper limit of
the studied interval. In the 10–20 kyr interval some d maximum between, the field strength seems to slowly increase to the upper limit of the studied interval. In the $10-20$ kyr interval some differences exist between individual records, and fine-scale details are not always on the other hand, well-resolved millennial-scale features are superimposed to the ual records, and fine-scale details are not always resolved. In the 20–75 kyr interval,
on the other hand, well-resolved millennial-scale features are superimposed to the
broader trends. The duration of the Laschamp event, on the other hand, well-resolved millennial-scale features are superimposed to the broader trends. The duration of the Laschamp event, which is recorded directionally in five cores, appears to be about 1500 years, consiste

in five cores, appears to be about 1500 years, consistent with a recent suggestion on the origin of geomagnetic excursions.

Keywords: palaeomagnetism; correlation; geomagnetic events; Atlantic Ocean

1. Introduction

1. Introduction
Study of geomagnetic field intensity recorded in marine sediments, in time and space,
has two principal objectives: firstly, to understand the origin of palaeointensity variahas two principal objectives: firstly, to understand the origin of palaeointensity varia-
has two principal objectives: firstly, to understand the origin of palaeointensity varia-
tions in terms of the geodynamo and second Study of geomagnetic field intensity recorded in marine sediments, in time and space,
has two principal objectives: firstly, to understand the origin of palaeointensity varia-
tions in terms of the geodynamo and, secondly Interior of the geodynamo and, secondly, to assess the potential of geomagnetic
intensity variations as a mean of stratigraphic correlation (see, for example, Peck *et* tions in terms of the geodynamo and, secondly, to assess the potential of geomagnetic
intensity variations as a mean of stratigraphic correlation (see, for example, Peck *et*
al. 1996). The Sint-200 (Guyodo & Valet 1996 intensity variations as a mean of stratigraphic correlation (see, for example, Peck *et* $al. 1996$). The Sint-200 (Guyodo & Valet 1996) and more recently the Sint-800 (Guyodo & Valet 1999) palaeointensity stacks, based re odo $\&$ Valet 1999) palaeointensity stacks, based respectively on 17 and 33 globally distributed palaeointensity records, have demonstrated the global synchronism of odo & Valet 1999) palaeointensity stacks, based respectively on 17 and 33 globally
distributed palaeointensity records, have demonstrated the global synchronism of
long wavelength $(10^4{-}10^5 \text{ kyr})$ features in the indiv distributed palaeointensity records, have demonstrated the global synchronism of long wavelength (10^4-10^5 kyr) features in the individual records. Shorter wavelength features are not apparent in these two stacks beca features are not apparent in these two stacks because of the low sedimentation rates
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Figure 1. Schematic map of the north Atlantic Ocean showing the location of the studied cores and the bathymetry (1000 m contour interval). The latitude and longitude of each site are reported with the water depths and the average sedimentation rate.

of individual records and the smoothing effect of stacking records with imperfect chronologies.

Recent coring of high deposition rate 'drift' deposits in the North Atlantic has led to the determination of a number of palaeointensity records from sediments with Recent coring of high deposition rate 'drift' deposits in the North Atlantic has
led to the determination of a number of palaeointensity records from sediments with
sedimentation rates in excess of 10 cm kyr⁻¹. Some aut led to the determination of a number of palaeointensity records from sediments with
sedimentation rates in excess of 10 cm kyr^{-1} . Some authors have suggested that
palaeointensity records may provide a powerful high-reso sedimentation rates in excess of 10 cm kyr⁻¹. Some authors have suggested that palaeointensity records may provide a powerful high-resolution method for global-
scale correlation (see, for example, Peck *et al.* 1996; St palaeointensity records may provide a powerful high-resolution method for global-
scale correlation (see, for example, Peck *et al.* 1996; Stoner *et al.* 1995); however,
some authors believe that variable amplitude and n scale correlation (see, for example, Peck *et al.* 1996; Stoner *et al.* 1995); however, some authors believe that variable amplitude and number of individual features limit the usefulness of palaeointensity records as a some authors believe that
the usefulness of palaeoi
(Schwartz *et al.* 1998).
Here we report new re e usefulness of palaeointensity records as a means of high-resolution correlation
chwartz *et al.* 1998).
Here we report new relative palaeointensity records for the 10–75 ka interval from
cNorth Atlantic cores separated

(Schwartz *et al.* 1998).
Here we report new relative palaeointensity records for the 10–75 ka interval from
six North Atlantic cores separated by as much as 5000 km, all of which are character-
ized by mean sedimentation Here we report new relative palaeointensity records for six North Atlantic cores separated by as much as 5000 km, ized by mean sedimentation rates of at least 10 cm kyr⁻¹. The produce a palaeointensity stack (NAPIS-75) or the 10–75 ka interval from
m, all of which are character-
These records are combined
0–75 ka period For this lim-**MATHEMATICAL,
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SCIENCES** six North Atlantic cores separated by as much as 5000 km, all of which are character-
ized by mean sedimentation rates of at least 10 cm kyr^{-1} . These records are combined
to produce a palaeointensity stack (NAPIS-75 ized by mean sedimentation rates of at least 10 cm kyr^{-1} . These records are combined
to produce a palaeointensity stack (NAPIS-75) for the $10-75$ ka period. For this lim-
ited time-interval, millennial-scale correlati to produce a palaeointensity stack (NAPIS-75) for the $10-75$ ka period. For this limited time-interval, millennial-scale correlation among cores with high sedimentation rate results in a very detailed relative palaeointe ited time-interval, millennial-scale correlation among cores with high sedimentation te results in a very detailed relative palaeointensity stack with well-defined short-
welength features.
Although NAPIS-75 is derived from the North Atlantic region only, it shows strik-
r similarity with other records fro

wavelength features.
Although NAPIS-75 is derived from the North Atlantic region only, it shows striking similarity with other records from the Blake Outer Ridge, with high-resolution
records from outside of the North Atla Although NAPIS-75 is derived from the North Atlantic region only, it shows striking similarity with other records from the Blake Outer Ridge, with high-resolution records from outside of the North Atlantic and with the lon ing similarity with other records from the Blake Outer Ridge, with high-resolution
records from outside of the North Atlantic and with the long-wavelength features
of Sint-200. Moreover, correlation with cosmogenic isotope records from outside of the North Atlantic and with the long-wavelength features
of Sint-200. Moreover, correlation with cosmogenic isotope production, as recorded
in the GRIP ice core, implies that the millennial-scale va of Sint-200. Moreover, correlation with cosmogenic isotope production, as recorded
in the GRIP ice core, implies that the millennial-scale variability is a fundamental
feature of the global scale geomagnetic field. This ge in the GRIP ice core, implies that the millennial-scale variability is a fundamental feature of the global scale geomagnetic field. This geomagnetic record, in addition, gives a precise estimate of the duration of the Lasc feature of the global scale geomagnetic field. This geomagnetic record, in addition,

2. Previous work

2. Previous work
The location of the cores is shown schematically in figure 1. Five cores were collected
between 58° N and 67° N and from 45° W to 4° E at different water denths, while the The location of the cores is shown schematically in figure 1. Five cores were collected
between 58° N and 67° N and from 45° W to 4° E at different water depths, while the
sixth was sampled at a lower latitude The location of the cores is shown schematically in figure 1. Five cores were collected
between 58° N and 67° N and from 45° W to 4° E at different water depths, while the
sixth was sampled at a lower latitude (33° N). Th between 58° N and 67° N and from 45° W to 4° E at different water depths, while the sixth was sampled at a lower latitude (33° N). They are all located in areas known for their high mean sedimentation rates (10–20 cm kyr their high mean sedimentation rates $(10-20 \text{ cm kyr}^{-1})$. The SU90 cores were collected *Phil. Trans. R. Soc. Lond.* A (2000)

during the PALEOCINAT-I cruise of the RV *Le Suroit* in 1990, while the MD95 cores
were recovered in 1995 during the IMAGES-I cruise of RV *Marion Dufresne* (Bassinot
& Labevrie 1996). Core PS2644-5 was obtained by the RV EERING during the PALEOCINAT-I cruise of the RV *Le Suroit* in 1990, while the MD95 cores
were recovered in 1995 during the IMAGES-I cruise of RV *Marion Dufresne* (Bassinot
& Labeyrie 1996). Core PS2644-5 was obtained by the RV during the PALEOCINAT-I cruise of the RV *Le Suroit* in 1990, while the MD95 cores were recovered in 1995 during the IMAGES-I cruise of RV *Marion Dufresne* (Bassinot & Labeyrie 1996). Core PS2644-5 was obtained by the RV *Polar Stern*. Site 983 was drilled during Ocean Drilling Program (ODP) Leg 162. A & Labeyrie 1996). Core PS2644-5 was obtained by the RV *Polar Stern*. Site 983 was drilled during Ocean Drilling Program (ODP) Leg 162. A description of these cores is given in Kissel *et al.* (1999, and references therein drilled during Ocean Drilling Program (ODP) Leg 162. A description of these cores
is given in Kissel *et al.* (1999, and references therein), Channell *et al.* (1997) and
Voelker *et al.* (1998).
Previous work has shown th is given in Kissel *et al.* (1999, and references therein), Channell *et al.* (1997) and

the different cores and is dominated by low-Ti-content magnetite with uniform grain sizes in the pseudo-single-domain range (Kissel *et al.* 1999). The bulk magnetic parameters of all these cores exhibit short-term variations which reflect small changes the different cores and is dominated by low-Ti-content magnetite with uniform grain sizes in the pseudo-single-domain range (Kissel *et al.* 1999). The bulk magnetic parameters of all these cores exhibit short-term variat sizes in the pseudo-single-domain range (Kissel *et al.* 1999). The bulk magnetic parameters of all these cores exhibit short-term variations which reflect small changes in the relative amount of the magnetic fraction wit parameters of all these cores exhibit short-term variations which reflect small changes
in the relative amount of the magnetic fraction within the detrital fraction, linked
to changes in the strength of the bottom current in the relative amount of the magnetic fraction within the detrital fraction, linked
to changes in the strength of the bottom current (Kissel *et al.* 1999). Previous work
on the ODP Site 983 has also established that the to changes in the strength of the bottom current (Kissel *et al.* 1999). Previous work
on the ODP Site 983 has also established that the magnetic mineralogy is suitable
for relative palaeointensity determinations over this the ODP Site 983 has also established that the magnetic mineralogy is suitable in relative palaeointensity determinations over this interval (Channell *et al.* 1997).
A precise millennial-scale correlation has been establ

for relative palaeointensity determinations over this interval (Channell *et al.* 1997).
A precise millennial-scale correlation has been established between five of these
cores for marine isotopic stage (MIS) 3 (27-60 ka) A precise millennial-scale correlation has been established between five of these cores for marine isotopic stage (MIS) 3 (27–60 ka) (Kissel *et al.* 1999). This correlation uses identification of detrital (Heinrich) laye cores for marine isotopic stage (MIS) 3 (27–60 ka) (Kissel *et al.* 1999). This corre-
lation uses identification of detrital (Heinrich) layers and Ash layer 2 in the first
step, then the record of short-lived oscillati anhysteretic remanent magnetization in the second step. Furthermore, using the step, then the record of short-lived oscillations in the magnetic susceptibility and anhysteretic remanent magnetization in the second step. Furthermore, using the correlation established by Voelker *et al.* (1998) betwee anhysteretic remanent magnetization in the second step. Furthermore, using the correlation established by Voelker *et al.* (1998) between the planktic δ^{18} O record of core PS2644-5 and the δ^{18} O record of the GISP correlation established by Voelker *et al.* (1998) between t
core PS2644-5 and the δ^{18} O record of the GISP2 ice core
all the data have been placed on the GISP2 age model. all the data have been placed on the GISP2 age model.

3. Methods

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The six cores have been subsampled continuously using *u*-channels (Tauxe *et al.*)

3. Methods
The six cores have been subsampled continuously using *u*-channels (Tauxe *et al.* 1983; Weeks *et al.* 1993). The natural remanent magnetization (NRM), the anhys-
teretic remanent magnetization (ARM) and the The six cores have been subsampled continuously using *u*-channels (Tauxe *et al.* 1983; Weeks *et al.* 1993). The natural remanent magnetization (NRM), the anhysteretic remanent magnetization (ARM) and the isothermal rem teretic remanent magnetization (ARM) and the isothermal remanent magnetization (IRM) were measured with a pass-through high-resolution DC-SQUIDs cryogenic teretic remanent magnetization (ARM) and the isothermal remanent magnetization (IRM) were measured with a pass-through high-resolution DC-SQUIDs cryogenic magnetometer in the shielded room of the LSCE. Stepwise in-line al (IRM) were measured with a pass-through high-resolution DC-SQUIDs cryogenic
magnetometer in the shielded room of the LSCE. Stepwise in-line alternating-field
(AF) demagnetization was used with an average of 8-10 steps. AR magnetometer in the shielded room of the LSCE. Stepwise in-line alternating-field (AF) demagnetization was used with an average of 8–10 steps. ARM was imparted along the axis of the u-channel using a 100 mT AF and $50 \mu T$ (AF) demagnetization was used with an average of 8–10 steps. ARM was imparted along the axis of the u-channel using a 100 mT AF and 50 μ T DC field. IRM was acquired along the Y-axis of the u-channel by passing it throu along the axis of the *u*-channel using a 100 mT AF and $50 \mu T$ DC field. IRM was
acquired along the *Y*-axis of the *u*-channel by passing it through the poles of an elec-
tromagnet. Low-field susceptibility was measured acquired along the Y-axis of the *u*-channel by passing it through the poles of an electromagnet. Low-field susceptibility was measured every centimetre on the *u*-channels using a Bartington loop sensor with a 45 mm diam tromagnet. Low-field susceptibility was measured every centimetre on the u -channels
using a Bartington loop sensor with a 45 mm diameter. Analyses of the hysteresis
parameters at 2.5–10 cm intervals downcore were perfor using a Bartington loop sensor with a 45 mm diameter. Analyses of the hysteresis
parameters at 2.5–10 cm intervals downcore were performed using an alternating gra-
dient force magnetometer (AGFM 2900). Thermomagnetic anal parameters at 2.5–10 cm intervals downcore w
dient force magnetometer (AGFM 2900). The
at a few representative depths in each core.

4. Results

(*a*) *Additional mineral magnetic studies*

For the five cores which had been previously examined only for the MIS3 (Kissel *et al.* 1999) we have extended the mineral magnetic studies through the MIS3 (Kissel et al. 1999) we have extended the mineral magnetic studies through the 10–75 ka interval. There are virtually no changes in the magnetic For the five cores which had been previously examined only for the MIS3 (Kissel *et al.* 1999) we have extended the mineral magnetic studies through the 10–75 ka interval. There are virtually no changes in the magnetic mi *et al.* 1999) we have extended the mineral magnetic studies through the 10–75 ka interval. There are virtually no changes in the magnetic mineralogy over this longer period with respect to the shorter interval considered interval. There are virtually no changes in the magnetic mineralogy over this longer
period with respect to the shorter interval considered by Kissel *et al.* (1999). For all
cores, the *S*-ratio (King & Channell 1991) va

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(b) \overline{C}
(b) 'Day' plots (Day *et al.* 1977).

(b) 'Day' plots (Day *et al.* 1977).

rapid excursions to values never lower than 0.85. This is consistent with a low-

coercivity magnetic mineralogy Curie-balance analyses of sediment samples from rapid excursions to values never lower than 0.85. This is consistent with a low-
coercivity magnetic mineralogy. Curie-balance analyses of sediment samples from
the different cores indicate that the magnetic mineralogy is rapid excursions to values never lower than 0.85. This is consistent with a low-
coercivity magnetic mineralogy. Curie-balance analyses of sediment samples from
the different cores indicate that the magnetic mineralogy is coercivity magnetic mineralogy. Curie-balance analyses of sediment samples from
the different cores indicate that the magnetic mineralogy is dominated by low-Ti
magnetite (figure 2). Isolated small black patches at the ver the different cores indicate that the magnetic mineralogy is dominated by low-Ti
magnetite (figure 2). Isolated small black patches at the very top of the studied
interval in core MD95-2034 (which could only be sampled fro magnetite (figure 2). Isolated small black patches at the very top of the studied
interval in core MD95-2034 (which could only be sampled from 15 m downward
because of coring disturbance in the upper part) indicate localiz

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filled and open circles correspond to projections onto the horizontal and vertical planes, respec-Figure 4. Representative AF demagnetization diagrams obtained from *u*-channel samples. The filled and open circles correspond to projections onto the horizontal and vertical planes, respectively. J_0 is the magnetizati filled and open circles correspond to projections on
tively. J_0 is the magnetization intensity before dem-
level (below sea floor) of the sample is indicated.

level (below sea floor) of the sample is indicated.
iron monosulphide (M. Cremer 1996, personal communication). An oxidized zone is iron monosulphide (M. Cremer 1996, personal communication). An oxidized zone is
present between 3.5 and 5 m depth at Site 983 (Channell *et al.* 1997) and between
32.5 and 35 m in core MD95-2034 iron monosulphide (M. Cremer 1996)
present between 3.5 and 5 m depth
32.5 and 35 m in core MD95-2034.
Plots of ARM versus volume susc 32.5 and 35 m in core MD95-2034.
Plots of ARM versus volume susceptibility (κ) are shown in figure 3a. In four of

the cores, tight grouping of the points along a line emanating from the origin of the Plots of ARM versus volume susceptibility (κ) are shown in figure 3*a*. In four of the cores, tight grouping of the points along a line emanating from the origin of the plot is consistent with very uniform magnetic gra the cores, tight grouping of the points along a line emanating from the origin of the plot is consistent with very uniform magnetic grain size (Banerjee *et al.* 1981; King *et al.* 1982, 1983). For core MD95-2034, the lin plot is consistent with very uniform magnetic grain size (Banerjee *et al.* 1981; King *et al.* 1982, 1983). For core MD95-2034, the line does not extrapolate through the origin, indicating that there is a significant par et al. 1982, 1983). For core MD95-2034, the line does not extrapolate through the origin, indicating that there is a significant paramagnetic contribution to the low field susceptibility. Hysteresis ratios, reported in the of figure 3b, tightly cluster within the pseudo-single domain (PSD) area, supporting field susceptibility. Hysteresis ratios, reported in the Day diagrams (Day *et al.* 1977) of figure 3*b*, tightly cluster within the pseudo-single domain (PSD) area, supporting the ARM- κ results. Changes in ARM and $\$ of figure 3b, tightly cluster within the pseudo-single domain (PSD) area, supporting
the ARM- κ results. Changes in ARM and κ also indicate that the concentration of
magnetic grains in the sediments varies by at mo the ARM- κ results. Changes in ARM and κ also indicate that the concentration of magnetic grains in the sediments varies by at most a factor of eight. These results indicate that the criteria for relative palaeoint magnetic grains in the sediments varies by at most a factor of eight. These results indicate that the criteria for relative palaeointensity estimates (see, for example, King *et al.* 1983) are satisfied in these North Atla

(b) *Extension of the correlation to the* $10-75$ *kyr period and to ODP Site 983*

Correlation between cores was carried out for the $10-75$ ka interval using an approach similar to the two-step approach used for the MIS3 interval by Kissel *et al.* (a) Extension of the corretation to the 10–75 kgr period and to ODP site 983
Correlation between cores was carried out for the 10–75 ka interval using an
approach similar to the two-step approach used for the MIS3 interva approach similar to the two-step approach used for the MIS3 interval by Kissel *et al.* (1999). Firstly, isotopic stage boundaries $5/4$, $4/3$ and $2/1$ and detrital (Heinrich) layer signals were used as tie points. Then (1999). Firstly, isotopic stage boundaries $5/4$, $4/3$ and $2/1$ and detrital (Heinrich) layer signals were used as tie points. Then, short-lived oscillations of the magnetic susceptibility were used to refine this corre layer signals were used as tie points. Then, short-lived oscillations of the magnetic
susceptibility were used to refine this correlation. These were, however, not quite as
clear over the entire interval as in MIS3. For co susceptibility were used to refine this correlation. These were, however, not quite as clear over the entire interval as in MIS3. For core MD95-2034 and ODP Site 983, the correlation could not be made for sediments younger clear over the entire interval as in MIS3. For core MD95-2034 and ODP Site 983, the correlation could not be made for sediments younger than 20 and 15 ka, respectively. For all other cores, 10 ka was the young age limit fo *Phil. Trans. R. Soc. Lond.* A (2000)

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Figure 5. Records of inclination derived from principal component analysis and NRM intensity after demagnetization at 25 or 30 mT versus age.

after demagnetization at 25 or 30 mT versus age.
All the cores have been correlated down to 75 ka and correlated to the GISP2 ice
core age allowing the GISP2 age model to be adopted All the cores have been correlated down to 75 ka and core age, allowing the GISP2 age model to be adopted. ore age, allowing the GISP2 age model to be adopted.
(*c*) *Natural remanent magnetization*

 (c) *Natural remanent magnetization*
NRM component directions were calculated every 5 cm by generating orthogonal
magnetization plots, some of which are shown in figure 4. A single stable compo-MRM component directions were calculated every 5 cm by generating orthogonal
demagnetization plots, some of which are shown in figure 4. A single stable compo-
nent of magnetization passing through the origin of the ortho NRM component directions were calculated every 5 cm by generating orthogonal
demagnetization plots, some of which are shown in figure 4. A single stable compo-
nent of magnetization passing through the origin of the ortho demagnetization plots, some of which are shown in figure 4. A single stable component of magnetization passing through the origin of the orthogonal projection was isolated at peak AF -fields of 20–25 mT using standard lea ment of magnetization passing through the origin of the orthogonal projection was
isolated at peak AF-fields of 20–25 mT using standard least-squares analysis. Maxi-
mum angular deviation (MAD) angles smaller than 5° isolated at peak AF-fields of $20-25$ mT using standard leas
mum angular deviation (MAD) angles smaller than 5° for the
indicate that NRM directions were precisely determined.
The inclinations derived from principal Im angular deviation (MAD) angles smaller than 5° for the great majority of cases
dicate that NRM directions were precisely determined.
The inclinations derived from principal component analysis (PCA) and NRM
censity

indicate that NRM directions were precisely determined.
The inclinations derived from principal component analysis (PCA) and NRM
intensity records of all the cores are given in figure 5. Inclinations fluctuate around
the v The inclinations derived from principal component analysis (PCA) and NRM
intensity records of all the cores are given in figure 5. Inclinations fluctuate around
the values expected for an axial geocentric dipole field at t intensity records of all the cores are given in figure 5. Inclinations fluctuate around
the values expected for an axial geocentric dipole field at the different sites, and with
a variability appropriate for geomagnetic se *Phil. Trans. R. Soc. Lond.* A (2000)

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Figure 6. (a) Component inclination records from the five cores which have recorded the Laschamp event and (b) comparison between component inclinations obtained using the u-channel technique and the component inclinations obtained from cubic samples (black dots) for core MD95-2034. Values of the maximum ang u -channel technique and the component inclinations obtained from cubic samples (black dots)

for core MD95-2034. Values of the maximum angular deviation (MAD) are also reported.
large inclination swing identified as the Laschamp event is recorded at *ca*. 40 ka.
As shown in the enlarged view of figure 6*a*, the sw large inclination swing identified as the Laschamp event is recorded at $ca. 40$ ka.
As shown in the enlarged view of figure $6a$, the swing has different amplitudes in
the different cores, probably reflecting different, b large inclination swing identified as the Laschamp event is recorded at ca 40 ka.
As shown in the enlarged view of figure $6a$, the swing has different amplitudes in
the different cores, probably reflecting different, bu As shown in the enlarged view of figure 6a, the swing has different amplitudes in
the different cores, probably reflecting different, but in all cases slight, degrees of
smoothing. In the two more northerly cores (PS2644the different cores, probably reflecting different, but in all cases slight, degrees of smoothing. In the two more northerly cores (PS2644-5 and SU90-24), negative inclinations with maximum values $ca. -45^{\circ}$ are observed nations with maximum values $ca. -45^{\circ}$ are observed. In the other two cores (SU90-33 and MD95-2009) very shallow positive inclinations are observed. Inclination reaches -90° for the southernmost record (MD95-2034), implying that smoothing is very slight in this core. This last result has been confirmed (figure $6b$) using standard -90° for the southernmost record (MD95-2034), implying that smoothing is very slight in this core. This last result has been confirmed (figure 6b) using standard $2 \times 2 \times 2$ cm³ discrete samples, to ascertain that slight in this core. This last result has been confirmed (figure 6b) using standard $2 \times 2 \times 2$ cm³ discrete samples, to ascertain that no perturbation has been introduced by the *u*-channel technique, in connection wit $2 \times 2 \times 2$ cm³ discrete samples, to ascertain that no perturbation has been introduced
by the *u*-channel technique, in connection with the large intensity drop related to
the Laschamp event (see Weeks *et al.* 1993). by the *u*-channel technique, in connection with the large intensity drop related to the Laschamp event (see Weeks *et al.* 1993). During this excursion, MAD values are similar to those observed outside this interval, as

The important result is that the directional changes of the Laschamp event are diagrams relative to a reverse direction (figure $4f$) and by figure $6b$.
The important result is that the directional changes of the Laschamp event are recorded over a very short time-interval. There are offsets in the The important result is that the directional changes of the Laschamp event are recorded over a very short time-interval. There are offsets in the timing of the event in the different cores but they are very slight and reac recorded over a very short time-interval. There are offsets in the timing of the event
in the different cores but they are very slight and reach a maximum of less than
500 years between cores SU90-24 and PS2644-5. They may in the different cores but they are very slight and reach a maximum of less than
500 years between cores SU90-24 and PS2644-5. They may be due either to small
errors in the correlation of the cores or, more realistically, 500 years between cores SU90-24 and PS2644-5. They may be due either to small
errors in the correlation of the cores or, more realistically, to slight differences in
the magnetization lock-in depth in the different cores. errors in the correlation of the cores or, more realistically, to slight differences in the magnetization lock-in depth in the different cores. These slight differences do not, however, significantly affect the determinati the magnetization lock-in depth in the different cores. These slight differences do
not, however, significantly affect the determination of the duration of the directional
changes of the Laschamp event which is about 1000 not, however, significantly affect the determination of the duration of the directional
changes of the Laschamp event which is about 1000 years in every case (figure 6).
The duration of the event (in direction) is estimate changes of the Laschamp event which is about 1000 years in The duration of the event (in direction) is estimated from the inclination lies outside the normal range of secular variation. Uniclination lies outside the normal range of secular variation.

(*d*) *Normalized record of magnetization*

Following Levi & Banerjee (1976) and King *et al*. (1982), ARM has been used Following Levi & Banerjee (1976) and King *et al.* (1982), ARM has been used
as the normalizer to obtain records of normalized remanence. Very similar results
were obtained using IRM or κ as normalizers (figure 7). Th Following Levi & Banerjee (1976) and King *et al.* (1982), ARM has been used
as the normalizer to obtain records of normalized remanence. Very similar results
were obtained using IRM or κ as normalizers (figure 7). Th as the normalizer to obtain records of normalized remanence. Very similar results
were obtained using IRM or κ as normalizers (figure 7). The normalized records were
generated using values of NRM and ARM after AF dema were obtained using IRM or κ as normalizers (figure 7). The normalized records were
generated using values of NRM and ARM after AF demagnetization at 25 mT, this
peak field being sufficient for complete removal of sec

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Figure 7. Comparison of palaeointensity profiles obtained using different normalizers
(ARM IRM and κ) for core PS2644-5 % of palaeointensity profiles obtained using (ARM, IRM and κ) for core PS2644-5.

Figure 8. Individual records of normalized field intensity versus age obtained for the six cores.

the NRM. For ODP Site 983, we have used the IRM normalized record given by the authors (Channell *et al*. 1997). e NRM. For ODP Site 983, we have used the IRM normalized record given by the thors (Channell *et al.* 1997).
Similarities among the six normalized records are illustrated in figure 8. Not ly the broad long-term characteri

authors (Channell *et al.* 1997).

Similarities among the six normalized records are illustrated in figure 8. Not

only the broad long-term characteristics, but also short-lived features, appear to be

recorded synchronou Similarities among the six normalized records are illustrated in figure 8. Not
only the broad long-term characteristics, but also short-lived features, appear to be
recorded synchronously in the six records. To construct t only the broad long-term characteristics, but also short-lived features, appear to be
recorded synchronously in the six records. To construct the stack, the six normalized
records were first interpolated to a common sampl recorded synchronously in the six records. To construct the stack, the six normalized
records were first interpolated to a common sampling interval of 100 years (corre-
sponding to *ca*. $1-2$ cm) to give equal weight t records were first interpolated to a common sampling interval of 100 years (corresponding to ca . 1–2 cm) to give equal weight to each record in the stack. They were then scaled to obtain a common palaeointensity value fo *Phil. Trans. R. Soc. Lond.* A (2000) **Phil.** *Phil. Trans. R. Soc. Lond.* A (2000)

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9. Stacked NAPIS-75 record obtained from the six studied cores. The shaded corresponds to the $\pm 2\sigma$ incertitude derived from the bootstrap calculation.

corresponds to the $\pm 2\sigma$ incertitude derived from the bootstrap calculation.
corresponding to the Laschamp event, and also to obtain a common average value
in the (10–75 kvr) interval. This technique has the advantage corresponding to the Laschamp event, and also to obtain a common average value
in the (10–75 kyr) interval. This technique has the advantage of not being overly
sensitive to abrupt variations in individual records. The sta corresponding to the Laschamp event, and also to obtain a common average value
in the (10–75 kyr) interval. This technique has the advantage of not being overly
sensitive to abrupt variations in individual records. The sta in the (10–75 kyr) interval. This technique has the advantage of not bei
sensitive to abrupt variations in individual records. The stack was then de
using the arithmetic mean at each common interpolated sampling point.
The sensitive to abrupt variations in individual records. The stack was then determined
using the arithmetic mean at each common interpolated sampling point.
The resulting stack is shown in figure 9 with associated uncertainti

using the arithmetic mean at each common interpolated sampling point.
The resulting stack is shown in figure 9 with associated uncertainties. Following
Tauxe *et al.* (1991) the confidence limits were assessed using a boot The resulting stack is shown in figure 9 with associated uncertainties. Following Tauxe *et al.* (1991) the confidence limits were assessed using a bootstrap approach: 'pseudo-samples' were obtained by randomly drawing se Tauxe *et al.* (1991) the confidence limits were assessed using a bootstrap approach:

'pseudo-samples' were obtained by randomly drawing series of six records from the

original records (each record may appear several ti 'pseudo-samples' were obtained by randomly drawing series of six recordinal records (each record may appear several times in a given pse Consistent with the recommendation of Hall (1988) we have used n^2 pse (here $n^2 =$ ² pseudo-samples (here $n^2 =$ records (each record may appear several times in a given pseudo-sample).
ent with the recommendation of Hall (1988) we have used n^2 pseudo-samples
 $2^2 = 6 \times 6 = 36$) to ensure that the distribution of mean records is co Consistent with the recommendation of Hall (1988) we have used n^2 pseudo-samples
(here $n^2 = 6 \times 6 = 36$) to ensure that the distribution of mean records is correctly
represented. The $\pm 2\sigma$ incertitude derived from t (here $n^2 = 6 \times 6 = 36$) to ensure that the distribution of mean records is correctly represented. The $\pm 2\sigma$ incertitude derived from this bootstrap calculation is shown as a grey shaded area in figure 9. In addition, w represented. The $\pm 2\sigma$ incertitude derived from this bootstrap calculation is shown
as a grey shaded area in figure 9. In addition, we have also used a jackknife approach
(see Caceci 1989) in which each individual reco as a grey shaded area in figure 9. In addition, we have also used a jackknife approach (see Caceci 1989) in which each individual record is in turn excluded from trial stacks. The departure of each trial stack from the mea (see Caceci 1989) in which each individual record is in turn excluded from trial stacks. The departure of each trial stack from the mean value of the stack yields the estimate of standard deviation. The incertitudes calcul stacks. The departure of each trial stack from the mean value of the stack yields the estimate of standard deviation. The incertitudes calculated using the two approaches are almost identical.
As a consequence of the high estimate of standard deviation. The incertitudes calculated using the two approaches

are almost identical.
As a consequence of the high degree of internal consistency between individual
records (figure 8), the six stacks of five cores obtained from the jackknife approach
are almost identical, and each indi As a consequence of the high degree of internal consistency between individual records (figure 8), the six stacks of five cores obtained from the jackknife approach are almost identical, and each individual record is very are almost identical, and each individual record is very similar to the stack itself (figure 10) except for a small difference observed for core SU90-33 in the 45-50 kyr
interval. High internal consistency is also indicated by values of the correlation coef-
inficients between the stack and individual rec interval. High internal consistency is also indicated by values of the correlation coefinterval. High internal consistency is also indicated by values of the correlation coef-
ficients between the stack and individual records, all of which are large and none
departs from the average value (0.76) by more tha ficients between the stack and individual records, all of which are large and none departs from the average value (0.76) by more than 2σ (table 1). The maximum discrepancy is observed for core SU90-33 which also has th discrepancy is observed for core SU90-33 which also has the lowest correlation coefficient (0.61) .

5. Discussion

5. Discussion
(*a*) *North Atlantic palaeointensity stack for the 10-75 ka interval (NAPIS-75)*

(a) North Atlantic palaeointensity stack for the $10-75$ ka interval (NAPIS-75)
Before assuming that the stacked record represents variations of the geomagnetic
field one must ascertain that it is not significantly affect Before assuming that the stacked record represents variations of the geomagnetic field, one must ascertain that it is not significantly affected by environmental factors.

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GISP2 age model (kyr)

GISP2 age model (kyr)
Figure 10. Stacked record NAPIS-75 superimposed on the individual records showing their
internal consistency 5 superimposed on the internal consistency. internal consistency.
Table 1. *Correlation coefficient between NAPIS-75 and the individual palaeointensity records*

cores	coeff. correl.
$NAPIS - PS2644-5$	0.832
$NAPIS - MD95-2009$	0.766
$NAPIS - SU90-33$	0.611
$NAPIS - ODP$ Site 983	0.812
$NAPIS - SU90-24$	0.731
$NAPIS - MD95-2034$	0.831

With this in mind, we have calculated the power (Blackman–Tukey) spectrum of the stacked record (figure 11*a*). Most of the power is distributed in a broad band With this in mind, we have calculated the power (Blackman–Tukey) spectrum of
the stacked record (figure 11a). Most of the power is distributed in a broad band
 ca 0.2 kyr⁻¹ ($T = 5$ kyr) where it exceeds a 95% confidence With this in mind, we have calculated the power (Blackman–Tukey) spectrum of
the stacked record (figure 11*a*). Most of the power is distributed in a broad band
 $ca. 0.2 \text{ kyr}^{-1}$ ($T = 5 \text{ kyr}$) where it exceeds a 95% conf the stacked record (figure 11*a*). Most of the power is distributed in a broad band $ca. 0.2 \,\text{kyr}^{-1}$ ($T = 5 \,\text{kyr}$) where it exceeds a 95% confidence level calculated from the first-order autoregressive process (red-no ca. 0.2 kyr⁻¹ ($T = 5$ kyr) where it exceeds a 95% confidence level calculated from
the first-order autoregressive process (red-noise generator). This spectrum is very
different from those of bulk parameters of individua the first-order autoregressive process (red-noise generator). This spectrum is very different from those of bulk parameters of individual cores, such as ARM, of the individual records which document a peak at *ca*. 0.7 ky different from those of bulk parameters of
individual records which document a peak
Dansgaard-Oeschger climatic oscillations.
We have also calculated the coherence dividual records which document a peak at $ca.0.7 \text{ kyr}^{-1}$ $(T = 1.4 \text{ kyr})$ related to ansgaard–Oeschger climatic oscillations.
We have also calculated the coherence in the frequency domain between the rmalized intensity an

Dansgaard–Oeschger climatic oscillations.
We have also calculated the coherence in the frequency domain between the
normalized intensity and the environmental sensitive susceptibility of the different cores (figure 11b). Coherence slightly above the 95% confidence level is occasionally observed, particularly at relatively high frequencies. Frequencies at which coherence

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spectral power (arbitrary units)

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Figure 11. (a) Power spectrum of the stacked record showing a significant broad power peak at Figure 11. (a) Power spectrum of the stacked record showing a significant broad power peak at frequencies ca . 0.2 kyr^{-1} . (b) Coherence of the stacked record with the low field susceptibility from each core. For this Figure 11. (a) Power spectrum of the stacked record showing a significant broad power peak at
frequencies ca. 0.2 kyr⁻¹. (b) Coherence of the stacked record with the low field susceptibility
from each core. For this cal frequencies *ca*. 0.2 kyr⁻¹. (*b*) Coherence of the stacked record with the low
from each core. For this calculation, the records have been smoothed to 50
coherence is observed *ca*. 0.2 kyr⁻¹. The 95% confidence leve

is observed differ slightly from core to core but in every case correspond to a unsignifis observed differ slightly from core to core but in every case correspond to a unsignificant power in the power spectrum. Conversely, no significant coherence is observed $ca \ 0.2 \ \text{kvr}^{-1}$ ($T = 5 \ \text{kvr}$) where the powe is observed differ slightly from core to core but in every case correspond to a unsignificant power in the power spectrum. Conversely, no significant coherence is observed $ca. 0.2 \,\text{kyr}^{-1}$ ($T = 5 \,\text{kyr}$) where the powe icant power in the power spectrum. Conversely, no significant coherence is observed $ca. 0.2 \,\text{kyr}^{-1}$ ($T = 5 \,\text{kyr}$) where the power of the stacked record exceeds the 95% confidence level. These results imply that the n ca. 0.2 kyr⁻¹ ($T = 5$ kyr) where the power of the stacked record exceeds the 95% con-
fidence level. These results imply that the normalized stack (figure 9) is largely free
of environmental/climatic influences and fait fidence level. These results imply that the normalized s
of environmental/climatic influences and faithfully des
of the geomagnetic field in the North Atlantic region.
NAPIS-75 documents that the intensity of the palaeon environmental/climatic influences and faithfully describes the relative variations
the geomagnetic field in the North Atlantic region.
NAPIS-75 documents that the intensity of the palaeomagnetic field has been highly
riabl

% of the geomagnetic field in the North Atlantic region.

NAPIS-75 documents that the intensity of the palaeomagnetic field has been highly

variable in the last 75 kyr. From 75 ka, the onset of the time-interval consider NAPIS-75 documents that the intensity of the palaeomagnetic field has been highly
variable in the last 75 kyr. From 75 ka, the onset of the time-interval considered
here, the field strength exhibits some oscillations, wit \blacktriangleright here, the field strength exhibits some oscillations, with a first minimum at *ca*. 65 ka, \blacktriangleright followed by a progressive increase to a broad maximum centred at *ca*. 50 ka. There is here, the field strength exhibits some oscillations, with a first minimum at $ca.65$ ka, followed by a progressive increase to a broad maximum centred at $ca.50$ ka. There is then an impressive drop to very low values, to a followed by a progressive increase to a broad maximum centred at ca . 50 ka. There is
then an impressive drop to very low values, to a minimum at 40 ka, corresponding to
the directional anomaly of the Laschamp event. The then an impressive drop to very low values, to a minimum at 40 ka, corresponding to
the directional anomaly of the Laschamp event. The duration of the palaeointensity
low, about 1500 years as measured by its half-width, is the directional anomaly of the Laschamp event. The duration of the palaeointensity
low, about 1500 years as measured by its half-width, is comparable with the estimated
duration of the directional change. After recovery, low, about 1500 years as measured by its half-width, is comparable with the estimated
duration of the directional change. After recovery, another intensity low observed
at *ca*. 34 ka (¹⁴C dated at *ca*. 28–30 ka corres duration of the directional change. After recovery, another intensity low observed
at *ca*. 34 ka $(^{14}C$ dated at *ca*. 28–30 ka corresponding to *ca*. 32–34 ka calendar age
(Voelker *et al.* 1998)) corresponds in age t at ca. 34 ka (¹⁴C dated at ca. 28–30 ka corresponding to ca. 32–34 ka calendar age (Voelker *et al.* 1998)) corresponds in age to the Mono Lake event (see Liddicoat 1992). After a high at 33 ka and two lows at 30 and 24 (Voelker *et al.* 1998)) corresponds in age to the Mono Lake event (see Liddicoat 1992). After a high at 33 ka and two lows at 30 and 24 ka with a broad maximum between, the field strength seems to slowly increase to the 1992). After a high at 33 ka and two lows at 30 and 24 ka with a broad maximum
between, the field strength seems to slowly increase to the upper limit of the studied
interval. Short-lived (millennial-scale) features are su between, the field str
interval. Short-lived
the entire interval. *Phil. Trans. R. Soc. Lond.* A (2000)

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Outer Ridge (Schwartz *et al*. 1998), Labrador Sea (Stoner *et al*. 1998), Lake Baikal (Peck *et* Outer Ridge (Schwartz *et al.* 1998), Labrador Sea (Stoner *et al.* 1998), Lake Baikal (Peck *et al.* 1996), Somali Basin (Meynadier *et al.* 1992), and Mediterranean Sea (Tric *et al.* 1992), and with Sint-200 (Guyodo & V al. 1996), Somali Basin (Meynadier et al. 1992), and Mediterranean Sea (Tric et al. 1992), and

(*b*) *Comparison with other records*

(b) Comparison with other records
Figure 12 compares NAPIS-75 with the Sint-200 composite record (Guyodo &
let 1996) and records obtained from the Labrador Sea (Stoner *et al.* 1998) the Valet 12 compares NAPIS-75 with the Sint-200 composite record (Guyodo & Valet 1996), and records obtained from the Labrador Sea (Stoner *et al.* 1998), the Blake Outer Ridge (Schwartz *et al.* 1998) the Lake Baikal (Peck Figure 12 compares NAPIS-75 with the Sint-200 composite record (Guyodo & Valet 1996), and records obtained from the Labrador Sea (Stoner *et al.* 1998), the Blake Outer Ridge (Schwartz *et al.* 1998), the Lake Baikal (Peck Valet 1996), and records obtained from the Labrador Sea (Stoner *et al.* 1998), the Blake Outer Ridge (Schwartz *et al.* 1998), the Lake Baikal (Peck *et al.* 1996), the Mediterranean (Tric *et al.* 1992) and the Somali Ba Mediterranean (Tric *et al.* 1992) and the Somali Basin (Meynadier *et al.* 1992). The long-wavelength trends of geomagnetic field intensity variations, such as the lows *ca.* 40 and 65 ka or the high *ca.* 50 ka, are pre long-wavelength trends of geomagnetic field intensity variations, such as the lows long-wavelength trends of geomagnetic field intensity variations, such as the lows $ca. 40$ and 65 ka or the high $ca. 50$ ka, are present in all these records, albeit sometimes at slightly different ages, presumably reflec ca. 40 and 65 ka or the high ca. 50 ka, are present in all these records, albeit sometimes
at slightly different ages, presumably reflecting inaccuracies in the chronologies of the
individual records and/or inconsistencie at slightly different ages, presumably reflecting inaccuracies in the chronologies of the individual records and/or inconsistencies between age models (NAPIS-75 is placed on the GISP2 age model, while all the other records individual records and/or inconsistencies between age models (NAPIS-75 is placed on the GISP2 age model, while all the other records use marine isotopic ages). However, NAPIS-75 does not exhibit the monotonic increase in intensity documented in the Sint-200 profile for the past 40 kyr. On the contrary, th NAPIS-75 does not exhibit the monotonic increase in intensity documented in the Sint-200 profile for the past 40 kyr. On the contrary, the intensity appears to have been rather stable in the 30–15 ka interval.
The record Sint-200 profile for the past 40 kyr. On the contrary, the intensity appears to have

the Blake Outer Ridge, is similar to NAPIS-75. Some very short-lived features of the

Figure 13. Stacked record NAPIS-75 (continuous line) and VADM record derived from the Figure 13. Stacked record NAPIS-75 (continuous line) and VADM record derived from the 36° Cl profile and expressed as VADM (dashed line). The latter is calculated on the basis of the relationship between the 36° C Figure 13. Stacked record NAPIS-75 (continuous line) and VADM record derived from the 36 Cl profile and expressed as VADM (dashed line). The latter is calculated on the basis of the relationship between the 36 Cl pr ³⁶Cl profile and expressed as VADM (dashed line). The latter is calculated on the basis of
the relationship between the ³⁶Cl production rate and the geomagnetic dipole field intensity
(Baumgartner *et al.* 1998). This the relationship between the ^{ov}Cl production rate and the geomagnetic dipole field intensity (Baumgartner *et al.* 1998). This curve was derived from the measurement of ³⁶Cl concentrations in the GRIP ice core. It has in the GRIP ice core. It has been low-pass filtered with a cut-off frequency of $1/3000 \text{ yr}^{-1}$ and then transferred to the GISP2 age model (Grootes & Stuiver 1997).

CH88-10P record, however, are not present in NAPIS-75, particularly in the $10{\text -}20$ ka CH88-10P record, however, are not present in NAPIS-75, particularly in the 10–20 ka
interval. Striking similarities are observed between NAPIS-75 and the records from
the Mediterranean, the Somali Basin and the Lake Baikal CH88-10P record, however, are not present in NAPIS-75, particularly in the 10–20 ka
interval. Striking similarities are observed between NAPIS-75 and the records from
the Mediterranean, the Somali Basin and the Lake Baikal interval. Striking similarities are observed between NAPIS-75 and the records from
the Mediterranean, the Somali Basin and the Lake Baikal, apart from the slightly
different age for the 65 ka low in the Somali Basin and L the Mediterranean, the Somali Basin and the Lake Baikal, apart from the slightly
different age for the 65 ka low in the Somali Basin and Lake Baikal records. Some
short-lived features are present in all four records (e.g. different age for the 65 ka low in the Somali Basin and Lake Baikal records. Some
short-lived features are present in all four records (e.g. in the 30–40 ka interval). The
Labrador Sea composite record is also similar to N short-lived features are present in all four records (e.g. in the 30–40 ka interval). The Labrador Sea composite record is also similar to NAPIS-75, both in long-term and short-lived features and its resolution appears to Labrador Sea composite record is also similar to NAPIS-75, both in long-term and
short-lived features and its resolution appears to be of the same order as NAPIS-75.
Some of the differences between the two records can be a short-lived features and its resolution appears to be of the same order as NAPIS-75.
Some of the differences between the two records can be attributed to chronologi-
cal discrepancies between the SPECMAP-derived chronology Some of the differences between the two records can be attributed to
cal discrepancies between the SPECMAP-derived chronology for the I
record (Stoner *et al.* 1998) and the GISP2 chronology for NAPIS-75. record (Stoner *et al.* 1998) and the GISP2 chronology for NAPIS-75.
(*c*) *Comparison with cosmogenic radionucleide production*

The strength of the geomagnetic field is the most important factor controlling The strength of the geomagnetic field is the most important factor controlling
cosmonucleide production. Because shielding of cosmic rays by the geomagnetic field
occurs at distances of several Earth radii from the surfac The strength of the geomagnetic field is the most important factor controlling
cosmonucleide production. Because shielding of cosmic rays by the geomagnetic field
occurs at distances of several Earth radii from the surface cosmonucleide production. Because shielding of cosmic rays by the geomagnetic field
occurs at distances of several Earth radii from the surface, only changes in the dipole
field affect shielding. Comparison of cosmonucleid occurs at distances of several Earth radii from the surface, only changes in the dipole
field affect shielding. Comparison of cosmonucleide flux records is a powerful way of
assessing the global (dipolar) nature of the var field affect shielding. Comparison of cosmonucleide flux records is a powerful way of assessing the global (dipolar) nature of the variability of palaeointensity records.
Previous work along these lines has shown that ¹⁰ assessing the global (dipolar) nature of the variability of palaeointensity records.
Previous work along these lines has shown that ¹⁰Be flux and geomagnetic field
intensity anticorrelate in a northern Atlantic core (Rob intensity anticorrelate in a northern Atlantic core (Robinson *et al.* 1995). In addition, a synthetic reconstruction of palaeofield intensity, inverted from a global stack intensity anticorrelate in a northern Atlantic core (Robinson *et al.* 1995). In addition, a synthetic reconstruction of palaeofield intensity, inverted from a global stack of ¹⁰Be marine deposition records, can be corr tion, a synthetic reconstruction of palaeofield intensity, inverted from a global stack
of ¹⁰Be marine deposition records, can be correlated to the SINT-200 palaeointensity
stack (Franck *et al.* 1997). More recently, i of ¹⁰Be marine deposition records, can be correlated to the SINT-200 palaeointensity
stack (Franck *et al.* 1997). More recently, it has been demonstrated that the 96–25 ka
record of ³⁶Cl flux from the GRIP ice core a stack (Franck *et al.* 1997). More recently, it has been demonstrated that the 96–25 ka
record of ³⁶Cl flux from the GRIP ice core agrees quite well with a production-rate
calculation based on a palaeointensity stack fr record of ³⁶Cl flux from the GRIP ice core agrees quite well with a production-rate calculation based on a palaeointensity stack from the Somali Basin (Baumgartner *et al.* 1998). In figure 13 we compare NAPIS-75 to the calculation based on a palaeointensity stack from the Somali Basin (Baumgartner *et al.* 1998). In figure 13 we compare NAPIS-75 to the synthetic field intensity record calculated from this same ³⁶Cl record in the 25-65 ations in the 36Cl same 36Cl record in the 25–65 ka interval, assuming that the variations in the ³⁶Cl flux are entirely due to modulation by the geomagnetic field. The ³⁶Cl-derived synthetic palaeointensity profile h calculated from this same 36 Cl record in the 25–65 ka interval, assuming that the variations in the 36 Cl flux are entirely due to modulation by the geomagnetic field. The 36 Cl-derived synthetic palaeointens ations in the ³⁶Cl flux are entirely due to modulation by the geomagnetic field. The ³⁶Cl-derived synthetic palaeointensity profile has been smoothed out using a 3000 year window in order to filter the solar modulation *Phil. Trans. R. Soc. Lond.* A (2000)

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Figure 14. Comparison of the NAPIS-75 record with: (a) the Mediterranean record (Tric *et al.* 1992); (b) the Somali Basin record (Meynadier *et al.* 1992); and (c) after adjusting the time-scales of the Somali and Mediter Figure 14. Comparison of the NAPIS-75 record with: (*a*) the Mediterranean record (Tric *et al.* 1992); (*b*) the Somali Basin record (Meynadier *et al.* 1992); and (*c*) after adjusting the time-scales of the Somali and stack.

stack.
age model, this profile has been transferred to the GISP2 age model using the results
of Grootes & Stuiver (1997). The striking similarities between the observed and synage model, this profile has been transferred to the GISP2 age model using the results of Grootes & Stuiver (1997). The striking similarities between the observed and syn-
thetic palaeointensity records (figure 13) implies age model, this profile has been transferred to the GISP2 age model using the results
of Grootes & Stuiver (1997). The striking similarities between the observed and syn-
thetic palaeointensity records (figure 13) implies of Grootes & Stuiver (1997). The striking similarities between the observed and synthetic palaeointensity records (figure 13) implies that the NAPIS-75 represents the geomagnetic field, with minimal environmental influenc the global-scale field, and therefore provides a template for global correlation. We geomagnetic field, with minimal environmental influence, and that NAPIS-75 reflects
the global-scale field, and therefore provides a template for global correlation. We
believe that the slight age offsets between NAPIS-75 the global-scale field, and therefore provides a template for global correlation. We
believe that the slight age offsets between NAPIS-75 and the ³⁶Cl-derived record can
be attributed to slight imperfections in the GISP be attributed to slight imperfections in the GISP2-GRIP or marine-GISP2 correlations. The coincidence of the maximum in 36 Cl flux (minimum synthetic palaeointensity) with the Laschamp event leaves little doubt that the increase in cosmogenic flux *ca*. 40 ka is due to low geomagnetic field stre tions. The coincidence of the maximum in 36 Cl flux (minimum synthetic palaeoin-
tensity) with the Laschamp event leaves little doubt that the increase in cosmogenic
flux *ca*. 40 ka is due to low geomagnetic field st tensity) with the Laschamp event leaves
flux *ca*. 40 ka is due to low geomagnetic
a supernova (see Sonnett *et al.* 1987).

(*d*) *High-resolution chronostratigraphic correlation based on palaeointensity records*

In figure 12, each palaeointensity record is placed on its own independent age In figure 12, each palaeointensity record is placed on its own independent age
model. Assuming the global nature of the NAPIS-75 stack, we adjust the chronolo-
gies of the Mediterranean and Somali Basin records by correlat In figure 12, each palaeointensity record is placed on its own independent age model. Assuming the global nature of the NAPIS-75 stack, we adjust the chronologies of the Mediterranean and Somali Basin records by correlatio model. Assuming the global nature of the NAPIS-75 stack, we adjust the chronologies of the Mediterranean and Somali Basin records by correlation of the palaeointensity records to NAPIS-75. We have chosen the most prominent gies of the Mediterranean and Somali Basin records by correlation of the palaeointensity records to NAPIS-75. We have chosen the most prominent features as initial tie points, then the small-scale features were adjusted by sity records to NAPIS-75. We have chosen the most prominent features as initial tie
points, then the small-scale features were adjusted by stretching and/or compressing
parts of the Mediterranean and Somali Basin records t points, then the small-scale features were adjusted by stretching and/or compressing
parts of the Mediterranean and Somali Basin records to maximize their correlation
coefficients with NAPIS-75. The maximum stretching/com coefficients with NAPIS-75. The maximum stretching/compression corresponds to a shift of about 5000 years from the original age model (this includes the difference between the marine age model and the GISP2 age model used here). After adjustshift of about 5000 years from the original age model (this includes the difference
between the marine age model and the GISP2 age model used here). After adjust-
ment, the three records can be superimposed (figure 14). Th between the marine age model and the GISP2 age model used here). After adjustment, the three records can be superimposed (figure 14). The correlation coefficients for the Mediterranean and Somali Basin records are 0.857 an ment, the three records can be superimposed (figure 14). The correlation coefficients
for the Mediterranean and Somali Basin records are 0.857 and 0.881, respectively,
similar to those obtained for the different cores used for the Mediterranean and Somali Basin records are 0.857 and 0.881, respectively, similar to those obtained for the different cores used for constructing NAPIS-75. Using palaeointensity records we have, therefore, achieved

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tion between North Atlantic and Somali Basin/Mediterranean cores as was obtained within the North Atlantic (to construct NAPIS-75) using more traditional correlation between North Atlantic and Somali Basin/Mediterranean cores as was obtained
within the North Atlantic (to construct NAPIS-75) using more traditional correla-
tion techniques. The correlation coefficients are also simi within the North Atlantic (to construct NAPIS-75) using more traditional correlation techniques. The correlation coefficients are also similar to those obtained by Peck *et al.* (1996) in the correlation of the Lake Baikal tion techniques. The correlation coefficients are also similar to those obtained by
Peck *et al.* (1996) in the correlation of the Lake Baikal palaeointensity stack with
the same Mediterranean and Somali Basin records. To Peck *et al.* (1996) in the correlation of the Lake Baikal palaeointensity stack with
the same Mediterranean and Somali Basin records. Together these results show that
millennial-scale variability of palaeointensity record the same Mediterranean and Somali Basin records. Together these results show that
millennial-scale variability of palaeointensity records can be used as a global-scale
correlation tool. However, it must be stressed that th millennial-scale variability of palaeointensity records can be used as a global-scale
correlation tool. However, it must be stressed that this correlation is certainly limited
to particular records. Of the five records whi correlation tool. However, it must be stressed that this correlation is certainly limited
to particular records. Of the five records which have been compared to NAPIS-75,
only two (the Mediterranean and Somali Basin record only two (the Mediterranean and Somali Basin records) covary with NAPIS-75 in fine-scale detail. For other palaeointensity records, either these features do not exist only two (the Mediterranean and Somali Basin records) covary with NAPIS-75 in
fine-scale detail. For other palaeointensity records, either these features do not exist
or it is impossible to make unequivocal correlation fro fine-scale detail. For other palaeointensity records, either these features do not exist
or it is impossible to make unequivocal correlation from core to core. The possibility
of obtaining this kind of correlation at mille or it is impossible to make unequivocal correlation from core to core. The possibility
of obtaining this kind of correlation at millennial scale must, therefore, be consid-
ered as exceptional and restricted to cores with of obtaining this kind
ered as exceptional and
magnetic mineralogy.
From a geomagnetic ed as exceptional and restricted to cores with high sedimentation rate and uniform
agnetic mineralogy.
From a geomagnetic point of view, the presence of coherent short-lived features
elds evidence that millennial-scale var

magnetic mineralogy.
From a geomagnetic point of view, the presence of coherent short-lived features
yields evidence that millennial-scale variability is a fundamental feature of the global-
scale (axial dipole) geomagneti From a geomagnetic point of view, t
yields evidence that millennial-scale vari
scale (axial dipole) geomagnetic field. (*e*) *Duration of the Laschamp event and nature of geomagnetic excursions*

The combination of different factors (precise intercorrelation of the cores, precise The combination of different factors (precise intercorrelation of the cores, precise
correlation to the GISP2 isotopic record, high sediment accumulation rate and uni-
form mineral magnetic properties) has permitted an est The combination of different factors (precise intercorrelation of the cores, precise
correlation to the GISP2 isotopic record, high sediment accumulation rate and uni-
form mineral magnetic properties) has permitted an es correlation to the GISP2 isotopic record, high sediment accumulation rate and uni-
form mineral magnetic properties) has permitted an estimate for the duration of the
Laschamp event (*ca*. 1500 years). This duration estima form mineral magnetic properties) has permitted an estimate for the duration of the Laschamp event $(ca.1500 \text{ years})$. This duration estimate is more precise than previous duration estimates for this (or any other) magnetic ex Laschamp event $(ca.1500 years)$. This duration estimate is more precise than previous duration estimates for this (or any other) magnetic excursion. It is important to note that the same value is obtained from each of the studi by vious duration estimates for this (or any other) magnetic excursion. It is important
to note that the same value is obtained from each of the studied cores, separated
by over 5000 km, so that it is not likely that the v to note that the same value is obtained from each of the studied cores, separated
by over 5000 km, so that it is not likely that the value is affected by local changes
in the sediment deposition rate at the very moment of by over 5000 km, so that it is not likely that the value is affected by local changes
in the sediment deposition rate at the very moment of the geomagnetic event. This
determination may have important implications in terms in the sediment deposition rate at the very moment of the geomagnetic event. This determination may have important implications in terms of processes in the Earth's
interior. Recently, Gubbins (1999) has proposed that excursions occur when the geo-
magnetic field reverses in the outer fluid core, which interior. Recently, Gubbins (1999) has proposed that excursions occur when the geo-
magnetic field reverses in the outer fluid core, which has a typical overturn time of
about 500 years, but not in the inner solid core, wh magnetic field reverses in the outer fluid core, which has a typical overturn time of
about 500 years, but not in the inner solid core, where diffusion of the field occurs on
a typical time-scale of 3000 years. The longer about 500 years, but not in the inner solid core, where diffusion of the field occurs on
a typical time-scale of 3000 years. The longer time constant of the inner core delays
full reversals of the field, during which time a typical time-scale of 3000 years. The longer time constant of the inner core delays full reversals of the field, during which time the original polarity may re-establish itself in the outer core, producing an excursion. full reversals of the field, during which time the original polarity may re-establish
itself in the outer core, producing an excursion. On the other hand, a full rever-
sal occurs when the field reverses in both the outer itself in the outer core, producing an excursion. On the other hand, a full reversal occurs when the field reverses in both the outer and inner cores, and requires a longer time than an excursion. The duration of the Lasch sal occurs when the field reverses in both the outer and inner cores, and requires a
longer time than an excursion. The duration of the Laschamp event determined from
NAPIS-75 is 1500 years at most when the GISP2 age model longer time than an excursion. The duration of the Laschamp event determined from NAPIS-75 is 1500 years at most when the GISP2 age model is used. If NAPIS-75 had been placed on a SPECMAP age model, the duration of the int NAPIS-75 is 1500 years at most when the GISP2 age model is used. If NAPIS-75 had been placed on a SPECMAP age model, the duration of the intensity low associated with the directional change would be somewhat longer, but in had been placed on a SPECMAP age model, the duration of the intensity low asso-
ciated with the directional change would be somewhat longer, but in both cases this
short duration is inconsistent with a full reversal occurr ciated with the directional change would be somewhat longer, but in both cases this
short duration is inconsistent with a full reversal occurring in the inner core. NAPIS-
75 therefore provides supporting evidence for Gubb short duration is inconsistent with a full reversal occurring in the inner core. NAPIS-
75 therefore provides supporting evidence for Gubbins's hypothesis of a distinction
between excursions and reversals, with the excursi 75 therefore provides support
between excursions and reve
occurring in the outer core.

6. Conclusions

6. Conclusions
We have obtained a precise record of geomagnetic palaeointensity from the north
Atlantic Ocean (NAPIS-75) characterized by high resolution and excellent time con-We have obtained a precise record of geomagnetic palaeointensity from the north
Atlantic Ocean (NAPIS-75), characterized by high resolution and excellent time con-Atlantic Ocean (NAPIS-75), characterized by high resolution and excellent time con-
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trol in the interval 75-10 ka. This record documents a picture of the geomagnetic field
with broad trends already recognized in other records and well-resolved short-lived trol in the interval 75–10 ka. This record documents a picture of the geomagnetic field
with broad trends already recognized in other records and well-resolved short-lived
features. By comparison with other palaeointensity *IATHEMATICAL,
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CIENCES* with broad trends already recognized in other records and well-resolved short-lived features. By comparison with other palaeointensity records and with the record of ³⁶Cl in the GRIP ice core we have established the global nature of the variability of Cl in the GRIP ice core we have established the global nature of the variability of
APIS-75.
A correlation between NAPIS-75 and published records of palaeointensity from
e Mediterranean and the Somali Basin obtained by mat NAPIS-75.

NAPIS-75.
A correlation between NAPIS-75 and published records of palaeointensity from
the Mediterranean and the Somali Basin obtained by matching long-term trends
and short-lived features of the palaeointensity records th A correlation between NAPIS-75 and published records of palaeointensity from
the Mediterranean and the Somali Basin obtained by matching long-term trends
and short-lived features of the palaeointensity records themselves i the Mediterranean and the Somali Basin obtained by matching long-term trends
and short-lived features of the palaeointensity records themselves is obtained with
the same accuracy as was achieved in the North Atlantic to co the same accuracy as was achieved in the North Atlantic to construct NAPIS-75 itself using more traditional correlation methods. This underlines the potential of the same accuracy as was achieved in the North Atlantic to construct
itself using more traditional correlation methods. This underlines the p
palaeointensity records for high-resolution correlation on a global scale.
The d elf using more traditional correlation methods. This underlines the potential of laeointensity records for high-resolution correlation on a global scale.
The duration of the Laschamp event, which is precisely recorded in f

The duration of the Laschamp event, which is precisely recorded in five of the cores both in direction and intensity, is about 1500–2000 years. This short duration The duration of the Laschamp event, which is precisely recorded in five of the cores both in direction and intensity, is about 1500–2000 years. This short duration provides evidence in favour of the hypothesis that excursi cores both in direction and intensity, is about 1500–2000 years. This short duration
provides evidence in favour of the hypothesis that excursions correspond to processes
in the outer core only, and have a shorter duration provides evidence in favour of the hypothesis that excursions core
in the outer core only, and have a shorter duration than reversal
the field reverses in both outer fluid core and inner solid core.

the field reverses in both outer fluid core and inner solid core.
We express our gratitude to L. Meynadier and J.-P. Valet, J. Peck, M. Schwartz and S. Lund,
and J. Stoner for providing us with the numerical data of their We express our gratitude to L. Meynadier and J.-P. Valet, J. Peck, M. Schwartz and S. Lund,
and J. Stoner for providing us with the numerical data of their palaeointensity records, and
to L. Beck, who carried out some of t and J. Stoner for providing us with the numerical data of their palaeointensity records, and
to L. Beck, who carried out some of the measurements for core MD95-2034. J. Stoner is also
thanked for his careful review of the to L. Beck, who carried out some of the measurements for core MD95-2034. J. Stoner is also to L. Beck, who carried out some of the measurements for core MD95-2034. J. Stoner is also
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